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# Prediction of Direct and Diffuse Solar Radiation on Pokhara, Nepal

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### Abstract

The measurement of direct and diffuse solar radiation is not easily possible because instruments are quite expensive. The main objective of this study was to predict direct solar radiation ( $I_B$ ) and diffuse solar radiation ( $I_D$ ) on Pokhara (28.186643° N, 83.97518° E, 800 m asl) for a period of one year (2017). Daily data of spectral Aerosol optical depth (AOD) and total ozone column are obtained from NASA website. The maximum value of direct solar radiation and diffuse solar radiation values were found 932.9  $\pm$  65.9 W/m² in January and 363.0  $\pm$  87.9 W/m² in February respectively. The minimum value of direct solar radiation and diffuse solar radiation values were found 692.8  $\pm$ 142.0 W/m² in April and 107.7  $\pm$  20.4 W/m² in July respectively. The diffuse solar radiation changes with season. The annual average of direct solar radiation and diffuse solar radiation are found 808.1 $\pm$  133.7 W/m² and 220.8  $\pm$  102.8 W/m² and respectively. The results of this research work will help further identification, impact and analysis of solar radiation in different locations in Nepal with the same geographic conditions.

Keywords: aerosol, atmospheric transmittance, diffuse, direct radiation, solar radiation

### 1. Introduction

Since the sun is the closest star to the earth, solar energy is our basic and primary source of energy. The Sun radiates  $4 \times 10^{26}$  W energy in forms of electromagnetic wave of wavelength  $0.3 \mu m$  (Poudyal et al., 2012). As the orbit of the Earth around the Sun is elliptical, the solar radiation decreases as square of distance between the Sun to the Earth.  $1367 \text{ W/m}^2$  ( $I_{sc}$ ) solar energy incidents on the outer surface of Earth's atmosphere when distance between Sun and Earth is  $1.49 \times 10^8$  km (Duffie & Beckman, 2013). The solar radiation ( $I_0$ ) incidents on specific point of outer surface of atmosphere at specific time depends on day number of year (DOY,  $n_d$ ) (Iqbal, 1983).

$$I_o = I_{sc} \left[ 1 + 0.033 \cos \left( \frac{2\pi}{365} n_d \right) \right]$$
 (1)

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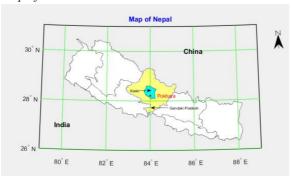
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When solar radiation passes through the atmosphere, it interacts with constituents of the atmosphere. The solar radiation is scattered, reflected, and absorbed in the atmosphere. The solar radiation is attenuated in atmosphere. The solar radiation incident on the ground in two ways, direct and diffuse (Liou, 2002). The sum of those two solar radiations; direct solar radiation ( $I_B$ ) and diffuse solar radiation ( $I_D$ ) gives global solar radiation (GSR) (Bason, 2004).

Nepal is situated between latitudes of 26.36° N to 30.45° N and longitudes of 80.06° E to 88.2° E in South East Asia. The elevation of the country ranges from 60 m to 8848 m within a span of 200 km from south to north and about 800 km from east to west (Majupuria,1999). In developing countries like Nepal, the most of energy consumption is fuel wood, agriculture residue, cow dung, coal and petroleum product. There is only 2.4 % (MoF, 2022) of energy consumption is alternative energy line hydroelectricity and solar. Nepal is land lock between two giant industrial countries India and China and their industrial byproduct can directly affects concentration of atmospheric components above Nepal. Nepal has to facility of sun tracker pyrheliometer and diffuse pyranometer to measure direct solar radiation and diffuse solar radiation respectively. Therefore, estimation of direct and diffuse solar radiation is necessary. Its application is very important on agriculture, Hydrology, Climate change, and energy harvesting (Poudyal et al., 2013).

Pokhara (28.186643° N, 83.97518° E, 800 m asl) lies in Kaski district of Gandaki Province as shown in Figure 1. Pokhara is capital of Gandaki Province and headquarter of Kaski district. Pokhara is the city of lakes (like Phewa, Begnas, Rupa). Three peaks Dhaulagiri, Annapurna, Manaslu with more than 8000 m height can be seen from the city. The Machhapuchchhre (Fishtail) with height of 6,993 metres is the closest to the city. Pokhara covers 464.24 square km. Pokhara has population 599,504 and population density 1,300 per square km (CBS, 2023). Average temperatures is between 25 and 35° C in summer and is -2 to 15° C in winter. Pokhara areas receive a high amount of precipitation in the country. Pokhara and nearby areas receive a high amount of precipitation 3,350 mm/year. Yearly mean daily GSR is 16.499 MJ/m²/day from 2007 to 2012 in Pokhara (Adhikari et al., 2013). Maximum GSR is 23.21 MJ/m²/day in June and minimum is 12.04 MJ/m²/day in December from 2009 to 2010 in Pokhara (Poudyal, 2015).

Figure 1
Map of Pokhara



Note. The map is sourced from the website of Survey Department, Government of Nepal

#### 2. Methods

The solar radiation is scattered, reflected, and absorbed by the atmospheric constituents like gas molecules, aerosols, water vapor, ozone and clouds. The solar radiation is attenuated in atmosphere with extinction coefficient (k) and optical air mass (m) (Lothian, 1963). The solar radiation incident on ground is two ways, direct and diffused.

Bird and Hulstrom (1981) developed Parameterization Model C for direct solar radiation (I<sub>B</sub>) and diffuse solar radiation (I<sub>D</sub>). Direct solar radiation (I<sub>B</sub>)

$$I_B = I_o(0.9751 \tau_o \tau_w \tau_o \tau_\alpha \tau_r + B(z)$$
 (2)

Here  $\tau_0$  is atmospheric transmittance due to ozone,  $\tau_w$  is atmospheric transmittance due to water vapor,  $\tau_g$  is atmospheric transmittance due to gas mixture,  $\tau_a$  is atmospheric transmittance due to aerosol and  $\tau_r$  is atmospheric transmittance due to Rayleigh scattering.  $\tau_0$  depends on ozone column (l) and relative air mass  $(m_r)$ .  $\tau_0$  depends on water content (w) and relative air mass.  $\tau_g$  and  $\tau_r$  depend absolute air mass  $(m_a)$ .  $\tau_a$  depends on Angstrom turbidity coefficient  $(\beta)$ , Angstrom exponential  $(\alpha)$  and absolute air mass (Wang et al., 2017). B(z) is a correction term for altitude (z) in meter (Bintanja,1996).

$$\begin{split} \tau_o &= 1 \, - \, [0.1611 \, u_3 (1 \, + \, 139.48 \, u_3)^{-0.3025} \, \, 0.002715 \, u_3 (1 \, + \, 0.044 \, u_3 \, + 0.0003 \, u_3^2)^{-1} \\ \tau_w &= 1 \, - \, 2.4959 \, \, u_1 [(1 \, + \, 79.034 \, u_1)^{0.682} \, + \, 6.385 \, u_1]^{-1} \\ \tau_g &= e^{[-0.0127 \, m_0^{0.26}]} \end{split}$$

$$\begin{split} \tau_{\alpha} &= (0.1244 \; \alpha \; - \; 0.0162) \; + \; (1.003 \; - \; 0.125 \; \alpha) e^{[-\beta \; m_{\alpha} (1.089 \; \alpha \; + \; 0.5123)]} \\ \tau_{\gamma} &= e^{[-0.0903 \; m_{\alpha}^{0.83} (1.01 \; + \; m_{\alpha} \; - \; m_{\alpha}^{1.01})]} \end{split}$$

Where

$$u_2 = l m_r$$

$$u_1 = w m_r$$

Diffuse solar radiation (I<sub>D</sub>) is

$$I_D = I_{dr} + I_{da} + I_{dm} \tag{3}$$

where  $I_{dr}$  is diffuse solar radiation produced by Rayleigh scattering,  $I_{da}$  is diffuse solar radiation produced by aerosols and  $I_{dm}$  is diffuse solar radiation produced by multiple reflections.

$$I_{dr} = 0.79I_0 cos\theta_z \tau_o \tau_g \tau_w \tau_{aa} \frac{0.5(1-\tau_r)}{1-m_v+m_v^{1.02}}$$
 (4)

Here  $\tau_{aa}$  is the atmospheric transmittance of aerosol absorptance.  $\theta z$  is solar zenith angle.

$$I_{dm} = \frac{(I_B \cos\theta_x + I_{dx} + I_{da}) \otimes_g \otimes_a}{1 - \otimes_g \otimes_a}$$

$$(5)$$

 $\alpha_g$  and  $\alpha_a$  are albedo of ground and atmospheric albedo respectively.

$$I_{d\alpha} = 0.79 I_0 cos\theta_z \ \tau_o \ \tau_g \ \tau_w \ \tau_{\alpha\alpha} \ F_c \frac{0.5(1-\tau_{\alpha\alpha})}{1-m_\alpha+m_\alpha^{1.02}} \eqno(6)$$

Here  $\tau_{as}$  is the atmospheric transmittance of aerosol scattering and  $F_c = 0.84$ .

Open source software Python 3.7 software is used to analysis data and to plot graph. Mean  $(\bar{x})$ , standard deviation  $(\sigma)$  are used as Statistical tool. Standard error (SE) is used as error bar in graph. Data is presented in form of  $\bar{x} \pm \sigma'$ . Quartile  $(Q_1, Q_2, Q_3)$ , skewness  $(\gamma_1)$  and  $(\gamma_2)$  are used to investigate nature of distribution of data.

$$SE = \frac{\sigma}{\sqrt{n}}$$
 (7)

Here, n is number of data

$$y_1 = \sqrt{\frac{\mu_3^2}{\mu_2^3}} \tag{8}$$

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$$y_2 = \frac{\mu_4}{\mu_2^2} - 3 \tag{9}$$

where  $\mu_2$ ,  $\mu_3$ ,  $\mu_4$  are  $2^{nd}$ ,  $3^{rd}$  and  $4^{th}$  moment respectively.

Correlation coefficient (r) is used to find relation between two variable x and y.

#### 3. Results

Daily ground base measurement of spectral aerosol optical depth (AOD) derive from AERONET of NASA and satellite measurement of Total ozone column (TOC) downloaded from NASA website for Pokhara for year 2017. Five wavelength ( $\lambda$ ) 675, 500, 440, 380 and 340 nm are used to calculate Angstrom turbidity coefficient ( $\beta$ ) and Angstrom exponential ( $\alpha$ ) by linear regression method for Angstrom model (Angstrom, 1961).

$$AOD = \beta \lambda^{-\alpha} \tag{10}$$

By using equation (1), (2), (3) and  $\tau_o$ ,  $\tau_w$ ,  $\tau_g$ ,  $\tau_a$ ,  $\tau_r$ , daily direct radiation (I<sub>B</sub>) and diffuse solar radiation (I<sub>D</sub>) are calculated. Daily diffuse solar radiation produced by Rayleigh scattering (I<sub>dr</sub>), diffuse solar radiation produced by aerosols (I<sub>da</sub>) and diffuse solar radiation produced by multiple reflections (I<sub>dm</sub>) are calculated by using equation (4), (5) and (6). Daily variation of I<sub>dr</sub>, I<sub>dm</sub> and I<sub>da</sub> are shown in Figure 2. Figure 2(a) shows daily variation of diffuse solar radiation produced by Rayleigh scattering (I<sub>dr</sub>). The maximum and minimum value of I<sub>dr</sub> during study period of 2017 are found 38.8 W/m<sup>2</sup> and 30.1 W/m<sup>2</sup> respectively. The annual mean and standard deviation are found 34.8 W/m<sup>2</sup> and 2.4 W/m<sup>2</sup> respectively. Figure 2(b) shows daily variation of diffuse solar radiation produced by multiple reflections (I<sub>dm</sub>). The maximum and minimum value of I<sub>dm</sub> during the study period are found 65.1 W/m<sup>2</sup> and 11.0 W/m<sup>2</sup> respectively. The annual mean and standard deviation are found 37.9 W/m<sup>2</sup> and 10.3 W/m<sup>2</sup> respectively. Figure 2(c) shows daily variation of diffuse solar radiation produced by aerosols (I<sub>da</sub>). The maximum and minimum value of I<sub>da</sub> during the study period are found 480.3 W/m<sup>2</sup> and 24.5 W/m<sup>2</sup> respectively. The annual mean and standard deviation are found 480.0 W/m<sup>2</sup> and 92.5 W/m<sup>2</sup> respectively.

Figure 2(d) shows daily variation of diffuse solar radiation ( $I_D$ ). The maximum and minimum value of  $I_D$  during the study period are found 580.4 W/m² on March 28 and 81.6 W/m² on July 22 respectively due to air pollution. The annual mean and standard deviation are found 220.8 W/m² and 102.8 W/m² respectively.

**Figure 2**Daily Variation of Diffuse Solar Radiation

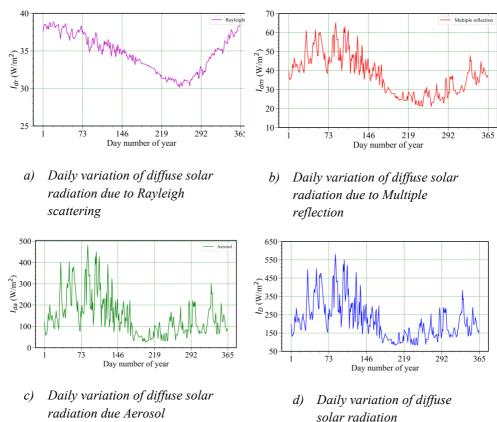
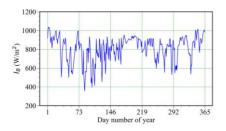
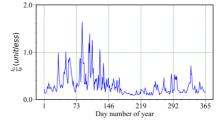


Figure 3(a) shows daily variation of direct solar radiation ( $I_B$ ) and figure 4(a) shows corresponding histogram. The maximum and minimum value of  $I_B$  during study period of 2017 are found 1037.1 W/m² on January 3 and 354.2 W/m² on March 28 respectively. The annual mean and standard deviation are found 808.1 W/m² and 133.7 W/m² respectively. The first quartile ( $Q_1$ ), second quartile ( $Q_2$ , median) and third quartile ( $Q_3$ ) are found 736.3 W/m², 841.0 W/m² and 910.1 W/m² respectively. Skewness ( $\gamma_1$ ) and kurtosis ( $\gamma_2$ ) are found -0.89 and 0.33 respectively. The distribution of  $I_B$  is negatively tailed due to negative value of skewness and is not normal (Gaussian). For normal distribution, kurtosis must be zero. Out of 313 days of study period, 102 days has direct solar radiation between 800 W/m² to 900 W/m². Figure 3(b) shows ratio of diffuse solar

radiation ( $I_D$ ) to direct solar radiation ( $I_B$ ). Out of 313 days of study period, 7 days have  $I_D$  is greater than  $I_B$ . Those days are cloudy day.

Figure 3
Daily Variation of Direct Solar Radiation

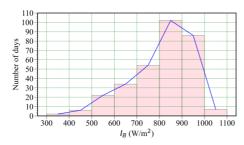


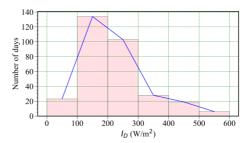


- a) Daily variation of direct solar radiation
- b) Ratio of diffuse solar radiation to direct solar radiation

Figure 4(b) shows histogram of diffuse radiation ( $I_D$ ). The first quartile ( $Q_1$ ), second quartile ( $Q_2$ , median) and third quartile ( $Q_3$ ) of diffuse solar radiation are found 144.1 W/m², 199.3 W/m² and 266.2 W/m² respectively. Skewness ( $\gamma_1$ ) and kurtosis ( $\gamma_2$ ) are found 1.12 and 0.94 respectively. The distribution of  $I_D$  is positively tailed as skewness is positive and is not Gaussian (normal). Out of 313 days of study period, 134 days has direct solar radiation between 100 W/m² to 200 W/m².

Figure 4
Histogram of Direct Solar Radiation and Diffuse Solar Radiation





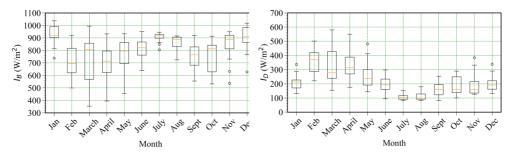
a) Histogram of direct solar radiation

b) Histogram of diffuse solar radiation

Figure 5 shows monthly variation of direct solar radiation ( $I_B$ ) and direct diffuse solar radiation ( $I_D$ ). Figure 5(a) shows monthly variation of direct solar radiation ( $I_B$ ). The maximum and minimum value of  $I_B$  during the study period are found 932.9  $\pm$  65.9

W/m² in January and  $692.8 \pm 142.0$  W/m² in April respectively. Variation is large in March due to large value of standard deviation 184.7 W/m² and less in July due less value in standard deviation 36.0 W/m². Figure 5(b) shows monthly variation of diffuse solar radiation ( $I_D$ ). The maximum and minimum value of  $I_D$  during the study period are found  $363.0 \pm 87.9$  W/m² in February and  $107.7 \pm 20.4$  W/m² in July respectively due clear and cloudy day. Variation is large in March due to large value of standard deviation 122.6 W/m² and less in July due less value in standard deviation 20.4W/m².

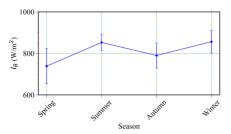
**Figure 5** *Monthly Variation of Direct Solar Radiation and Diffuse Solar Radiation* 

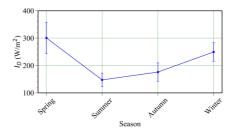


- a) Monthly variation of direct solar radiation
- b) Monthly variation of diffuse solar radiation

Figure 6 shows seasonal variation of direct solar radiation ( $I_B$ ) and direct diffuse solar radiation ( $I_D$ ). Figure 6(a) shows seasonal variation of direct solar radiation ( $I_B$ ). The maximum and minimum value of  $I_B$  during the study period are found 856.2  $\pm 93.2$  W/m² in Winter and 738.5  $\pm$  146.5 W/m² in spring respectively due to air pollution. Variation is large in spring due to large value of standard deviation 146.5 W/m² and less in summer due less value in standard deviation 66.8 W/m². Figure 6(b) shows seasonal variation of diffuse solar radiation ( $I_D$ ). The maximum and minimum value of  $I_D$  during the study period are found 300.7  $\pm$  99.6 W/m² in spring and 147.3  $\pm$  43.1 W/m² in summer respectively. Variation is large in spring due to large value of standard deviation 99.6 W/m² and less in summer due less value in standard deviation 43.1 W/m².

**Figure 6**Seasonal Variation of Direct Solar Radiation and Diffuse Solar Radiation





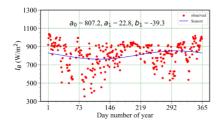
- a) Seasonal variation of direct solar radiation
- b) Seasonal variation of diffuse solar radiation

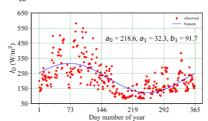
Seasonal variation of time series data are also analyzed by using Fourier series (Anton et al., 2011).

$$y_s = a_0 + a_1 \cos(\frac{2\pi}{365}n_d) + b_1 \sin(\frac{2\pi}{365}n_d)$$
 (11)

Here  $a_0$  is offset and  $\sqrt{a_1^2 + b_1^2}$  gives amplitude of seasonal component. Offset and amplitude of seasonal component of direct solar radiation are found 807.2 W/m<sup>2</sup> and 45.4 W/m<sup>2</sup> respectively as shown in Figure 7(a). Offset and amplitude of seasonal component of diffuse solar radiation are found 218.6 W/m<sup>2</sup> and 97.3 W/m<sup>2</sup> respectively as shown in Figure 7(b). Due to large value of amplitude of seasonal component, diffuse solar radiation is seasonal dependent.

**Figure 7**Fourier Analysis of Direct Solar Radiation and Diffuse Solar Radiation

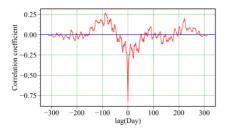


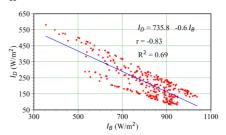


- a) Fourier analysis of direct solar radiation
- b) Fourier analysis of diffuse solar radiation

Figure 8(a) shows cross correlation between  $I_D$  and  $I_B$  with time lag in day. Correlation coefficient (r) is -0.83 at zero-day time lag. They are negatively corrected. When  $I_D$  increases,  $I_B$  decreases. They are fitted in straight line (y = mx + c) by best fit method as shown in figure 8(b). The straight line obtained with slope -0.6, y intercept is 735.8 and 69 % coefficient of determination ( $\mathbb{R}^2$ ).

Figure 8
Variation between Direct Solar Radiation and Diffuse Solar Radiation





- a) Cross correlation between  $I_B$  and  $I_D$
- b) Straight line

# 4. Discussion

Datasets on solar radiation is of great importance to the detection of global dimming and brightening. Direct solar radiation and diffuse solar radiation estimation using easily available satellite data on basis of atmospheric transmittance. They are therefore important to propose solar energy potential of the location, the sustainable development of ecological environments and agriculture-based productivity. In study period of one year (2017) for Pokhara, the annual average of direct solar radiation and diffuse solar radiation are found  $808.1\pm133.7~\text{W/m}^2$  and  $220.8\pm102.8~\text{W/m}^2$  respectively at 12 PM NST. The Annual average of direct solar radiation is found  $1106\pm70~\text{W/m}^2$  in Jomsom for a year 2012 (Shrestha et al., 2022). The annual average of diffuse solar radiation and direct solar radiation are found  $216\pm88~\text{W/m}^2$  and  $216\pm88~\text{W/m}^2$  respectively on Bode, Bhaktapur on 2013 (Shrestha et al., 2022a). The annual average of diffuse solar radiation on Beijing and Lasa of China were to be 6.6 MJ/m²/day (7.6 W/m²) and 6.1 MJ/m²/day (7.1 W/m²) respectively from 1993 to 2015(Wang et al., 2019). Diffuse solar radiation is large on Pokhara than big city of China and India due to large amount of precipitation and air pollution.

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